

Unexpected features of non-locality: resonant particle production



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Based on...

Resonant particle creation by a time-dependent potential in a nonlocal theory

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Considering an exactly solvable local quantum theory of a scalar field interacting with a δ -shaped time-dependent potential we calculate the Bogoliubov coefficients analytically and determine the spectrum of created particles. We then show how these considerations, when suitably generalized to a specific nonlocal “infinite-derivative” quantum theory, are impacted by the presence of nonlocality. In this model, nonlocality leads to a significant resonant amplification of certain modes, leaving its imprint not only in the particle spectrum but also in the total number density of created particles.

arXiv:2011.12929 [hep-th], submitted to Physics Letters B.

More work on non-local theories with Valeri Frolov and Andrei Zelnikov.

Motivation

Non-locality plays an important role in modern physics:

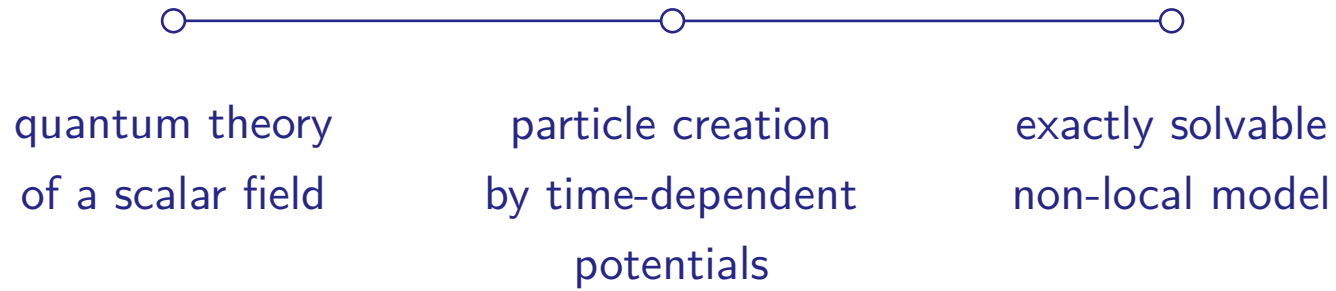
- Entanglement is a non-local phenomenon in quantum theory.
- Many effective actions in quantum field theory are non-local.
- It is probably impossible to define local observables in quantum gravity.
- Non-locality may solve the black hole information loss problem.

But: non-locality also challenges many of our “standard” notions in theoretical physics.

- Causality is typically violated at some scale (and perhaps beyond).
- Variational principle is not necessarily self-consistent.
- Notion of a “local particle” is difficult to define.

Solution: discuss reasonably well-behaved non-local theories, and use asymptotic local properties.

Unexpected features of non-locality: resonant particle production



PART I

QUANTUM THEORY OF A SCALAR FIELD

creation and annihilation operators ■ definition of vacuum

Classical scalar field

A free classical scalar field in Minkowski spacetime:

$$\square\varphi(t, \mathbf{x}) = 0, \quad \square = -\partial_t^2 + \nabla^2$$

We don't know any real scalar fields in Nature, use it as a toy model.

Use a Fourier transform to express a solution of this field equation:

$$\varphi(t, \mathbf{x}) = \int \frac{d^d k}{(2\pi)^{d/2}} \frac{1}{\sqrt{2\omega_{\mathbf{k}}}} \left[e^{-i\omega_{\mathbf{k}}t + i\mathbf{k}\cdot\mathbf{x}} a_{\mathbf{k}} + e^{+i\omega_{\mathbf{k}}t - i\mathbf{k}\cdot\mathbf{x}} a_{\mathbf{k}}^* \right]$$

The frequency $\omega_{\mathbf{k}}$ is given by the dispersion relation $\omega_{\mathbf{k}}^2 = \mathbf{k}^2$.

The coefficients $a_{\mathbf{k}}$ are the (complex) Fourier components of the field φ and satisfy $a_{-\mathbf{k}} = a_{\mathbf{k}}^*$.

So how can we make this a quantum theory?

Quantum scalar field

A free **quantum** scalar field in Minkowski spacetime:

$$\square \hat{\phi}(t, \mathbf{x}) = 0, \quad \square = -\partial_t^2 + \nabla^2$$

This equation is a differential equation for a “field operator.”

Use a Fourier transform to express a solution of this field equation:

$$\hat{\phi}(t, \mathbf{x}) = \int \frac{d^d k}{(2\pi)^{d/2}} \frac{1}{\sqrt{2\omega_{\mathbf{k}}}} \left[e^{-i\omega_{\mathbf{k}}t + i\mathbf{k} \cdot \mathbf{x}} \hat{a}_{\mathbf{k}} + e^{+i\omega_{\mathbf{k}}t - i\mathbf{k} \cdot \mathbf{x}} \hat{a}_{\mathbf{k}}^\dagger \right]$$

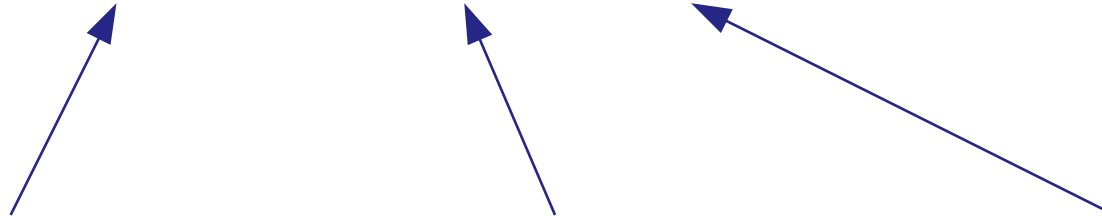
The frequency $\omega_{\mathbf{k}}$ is given by the dispersion relation $\omega_{\mathbf{k}}^2 = \mathbf{k}^2$.

The operators $\hat{a}_{\mathbf{k}}^\dagger$ and $\hat{a}_{\mathbf{k}}$ are so-called creation and annihilation operators of the field $\hat{\phi}$.

They can act on the vacuum state $|0\rangle$ to create or remove quanta of momentum \mathbf{k} .

What is the main idea of these operators?

$$|\Psi\rangle = |m_{\mathbf{k}}, n_{\mathbf{q}}, \dots\rangle$$



state of the
entire system

number of particles
with momentum \mathbf{k} ,
add one with $\hat{a}_{\mathbf{k}}^\dagger$,
remove one with $\hat{a}_{\mathbf{k}}$.

number of particles
with momentum \mathbf{q} ,
add one with $\hat{a}_{\mathbf{q}}^\dagger$,
remove one with $\hat{a}_{\mathbf{q}}$.

Some more basics on creation and annihilation operators

Definition of the vacuum state:

$$\hat{a}_{\mathbf{k}}|0\rangle = 0$$

Commutation relations (needed for canonical commutation relations at the Hamiltonian level):

$$[\hat{a}_{\mathbf{k}}, \hat{a}_{\mathbf{k}'}^\dagger] = \delta^{(d)}(\mathbf{k} - \mathbf{k}'), \quad [\hat{a}_{\mathbf{k}}, \hat{a}_{\mathbf{k}'}] = [\hat{a}_{\mathbf{k}}^\dagger, \hat{a}_{\mathbf{k}'}^\dagger] = 0$$

The concrete representation of $\hat{a}_{\mathbf{k}}^\dagger$ and $\hat{a}_{\mathbf{k}}$ is not so important. Use basic relations:

$$\hat{a}_{\mathbf{k}}^\dagger |n_{\mathbf{k}}\rangle = \sqrt{n_{\mathbf{k}} + 1} |n_{\mathbf{k}} + 1\rangle, \quad \hat{a}_{\mathbf{k}} |n_{\mathbf{k}}\rangle = \sqrt{n_{\mathbf{k}}} |n_{\mathbf{k}} - 1\rangle$$

Then, many-particle states can be written like this:

$$|m_{\mathbf{k}}, n_{\mathbf{q}}, \dots\rangle \sim (\hat{a}_{\mathbf{k}}^\dagger)^m |0\rangle \otimes (\hat{a}_{\mathbf{q}}^\dagger)^n |0\rangle \otimes \dots$$

You can count particles with the occupation number density operator $\hat{n}_{\mathbf{k}} = \hat{a}_{\mathbf{k}}^\dagger \hat{a}_{\mathbf{k}}$.



quantum theory
of a scalar field

particle creation
by time-dependent
potentials

exactly solvable
non-local model



PART II

PARTICLE CREATION BY POTENTIALS

potentials ■ Bogoliubov coefficients ■ Lippmann–Schwinger equation

Quantum scalar field with time-dependent potential

A quantum scalar field in Minkowski spacetime with **time-dependent potential**:

$$[\square - V(t)] \hat{\varphi}(t, \mathbf{x}) = 0, \quad \square = -\partial_t^2 + \nabla^2$$

This equation is again a differential equation for the “field operator” $\hat{\varphi}(t, \mathbf{x})$.

Use a purely spatial Fourier transform to rewrite a solution of this field equation:

$$\hat{\varphi}(t, \mathbf{x}) = \int \frac{d^d k}{(2\pi)^{d/2}} \left[e^{+i\mathbf{k} \cdot \mathbf{x}} \hat{\varphi}_{\mathbf{k}}(t) + e^{-i\mathbf{k} \cdot \mathbf{x}} \hat{\varphi}_{\mathbf{k}}^\dagger(t) \right]$$

Can now study the equation for the operators $\hat{\varphi}_{\mathbf{k}}(t)$ directly:

$$\left[\partial_t^2 + \omega_{\mathbf{k}}^2 + V(t) \right] \hat{\varphi}_{\mathbf{k}}(t) = 0$$

This is still a very complicated problem.

Quantum scalar field with time-dependent potential

Let's assume that the potential vanishes at early and late times, $(\partial_t^2 + \omega_{\mathbf{k}}^2) \hat{\varphi}_{\mathbf{k}}(t) = 0$.

At those “asymptotic times” there is no potential term and have again the free solution from before:

$$\hat{\varphi}_{\mathbf{k}}(t \rightarrow -\infty) = \frac{e^{-i\omega_{\mathbf{k}}t}}{\sqrt{2\omega_{\mathbf{k}}}} \hat{a}_{\mathbf{k}}, \quad \hat{\varphi}_{\mathbf{k}}(t \rightarrow +\infty) = \frac{e^{-i\omega_{\mathbf{k}}t}}{\sqrt{2\omega_{\mathbf{k}}}} \hat{b}_{\mathbf{k}} + \frac{e^{+i\omega_{\mathbf{k}}t}}{\sqrt{2\omega_{\mathbf{k}}}} \hat{b}_{\mathbf{k}}^\dagger$$

Note the different creation and annihilation operators at **early times** and at **late times**.

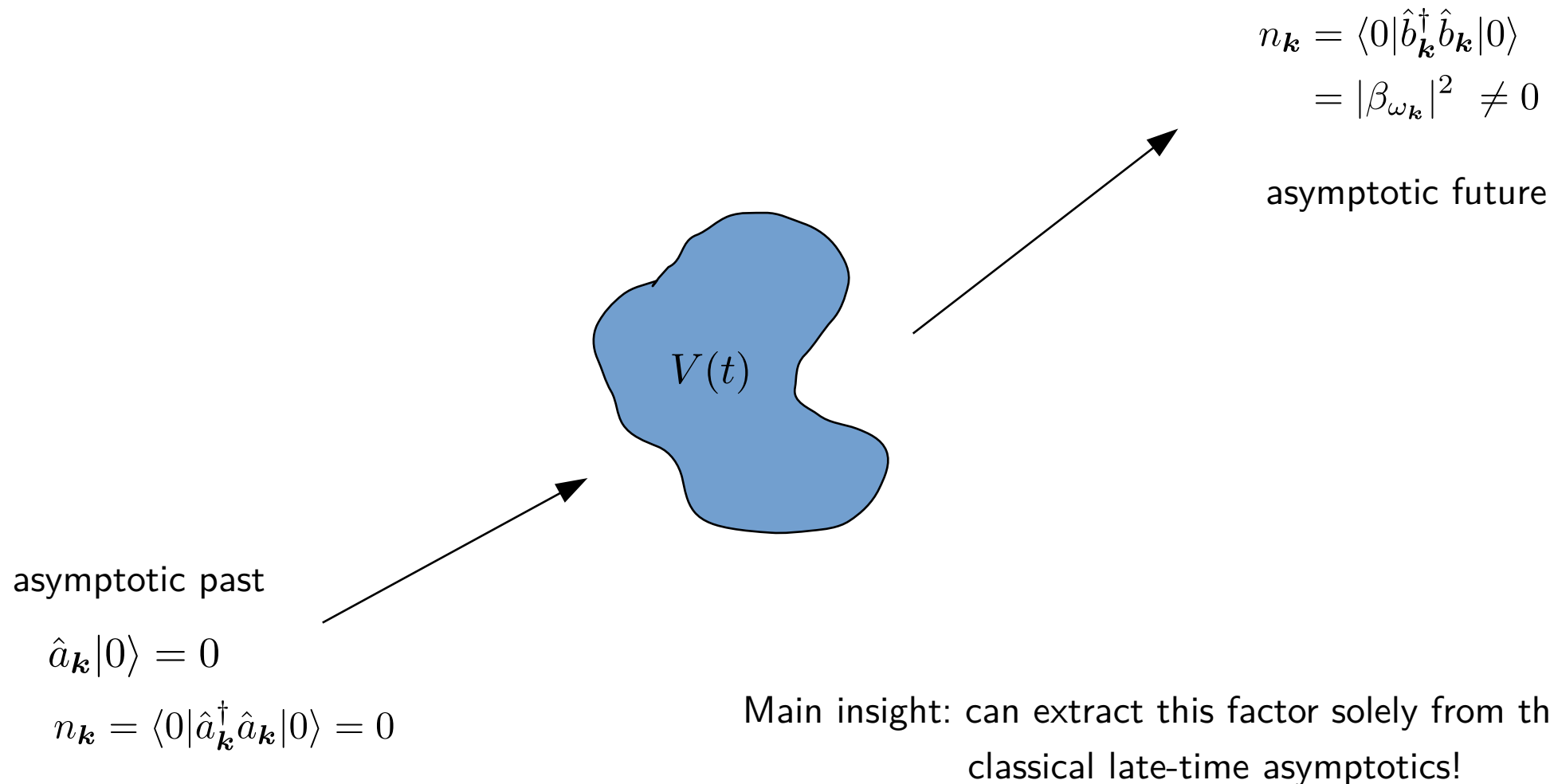
There are two different vacua (“in”-vacuum and “out”-vacuum):

$$\hat{a}_{\mathbf{k}}|\text{in}; 0\rangle = 0, \quad \hat{b}_{\mathbf{k}}|\text{out}; 0\rangle = 0$$

Bogoliubov coefficients relate these operators ($|\alpha_{\omega_{\mathbf{k}}}|^2 - |\beta_{\omega_{\mathbf{k}}}|^2 = 1$ ensures canonical trafo):

$$\hat{b}_{\mathbf{k}} = \alpha_{\omega_{\mathbf{k}}} \hat{a}_{\mathbf{k}} + \beta_{\omega_{\mathbf{k}}}^* \hat{a}_{-\mathbf{k}}^\dagger, \quad \hat{b}_{\mathbf{k}}^\dagger = \alpha_{\omega_{\mathbf{k}}}^* \hat{a}_{\mathbf{k}}^\dagger + \beta_{\omega_{\mathbf{k}}} \hat{a}_{-\mathbf{k}}$$

And what does all of this have to do with particle creation?

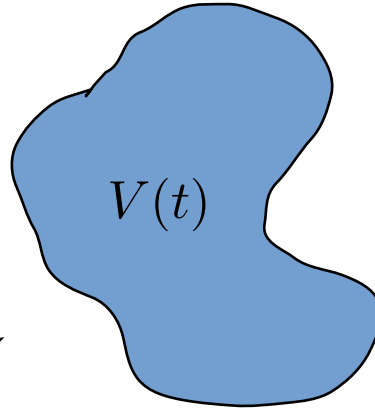


And what does all of this have to do with particle creation?



$$\hat{a}_{\mathbf{k}}|0\rangle = 0$$

$$n_{\mathbf{k}} = \langle 0 | \hat{a}_{\mathbf{k}}^\dagger \hat{a}_{\mathbf{k}} | 0 \rangle = 0$$



$$\begin{aligned} n_{\mathbf{k}} &= \langle 0 | \hat{b}_{\mathbf{k}}^\dagger \hat{b}_{\mathbf{k}} | 0 \rangle \\ &= |\beta_{\omega_{\mathbf{k}}}|^2 \neq 0 \end{aligned}$$



Main insight: can extract this factor solely from the classical late-time asymptotics!

Lippmann–Schwinger equation

How to extract the late-time asymptotics? Use the retarded Green function:

$$(\partial_t^2 + \omega_{\mathbf{k}}^2)\varphi_{\mathbf{k}}(t) = -V(t)\varphi_{\mathbf{k}}(t)$$

$$(\partial_t^2 + \omega_{\mathbf{k}}^2)G_{\mathbf{k}}^{\text{R}}(t' - t) = -\delta(t' - t), \quad G_{\mathbf{k}}^{\text{R}}(t' - t) = 0 \text{ if } t' < t.$$

Then, the solution can be written in Lippmann–Schwinger form like this:

$$\varphi_{\mathbf{k}}(t) = \varphi_{\mathbf{k}}^0(t) + \int_{-\infty}^{\infty} dt' G_{\mathbf{k}}^{\text{R}}(t - t') V(t') \varphi_{\mathbf{k}}(t')$$

Important: $\varphi_{\mathbf{k}}^0$ is a free solution that encodes the asymptotic past (because $G_{\mathbf{k}}^{\text{R}}$ vanishes there).

$$\varphi_{\mathbf{k}}^0(t) = \frac{1}{\sqrt{2\omega_{\mathbf{k}}}} e^{-i\omega_{\mathbf{k}} t}$$

Lippmann–Schwinger equation

This integral collapses for a delta-shaped potential $V(t) = \lambda\delta(t)$.

This potential also vanishes at early and late times, so we can expect a reasonable result.

$$\varphi_{\mathbf{k}}(t) = \varphi_{\mathbf{k}}^0(t) + \int_{-\infty}^{\infty} dt' G_{\mathbf{k}}^{\text{R}}(t - t') V(t') \varphi_{\mathbf{k}}(t') = \varphi_{\mathbf{k}}^0(t) + \lambda G_{\mathbf{k}}^{\text{R}}(t) \varphi_{\mathbf{k}}(0)$$

$$\varphi_{\mathbf{k}}(0) = \frac{\varphi_{\mathbf{k}}^0(0)}{1 - \lambda G_{\mathbf{k}}^{\text{R}}(0)}$$

The local retarded Green function has this form (sum of positive and negative frequencies):

$$G_{\mathbf{k}}^{\text{R}}(t' - t) = \frac{i}{2\omega_{\mathbf{k}}} \left[e^{+i\omega_{\mathbf{k}}(t' - t)} - e^{-i\omega_{\mathbf{k}}(t' - t)} \right] \theta(t' - t)$$

Extracting the Bogoliubov coefficient

Collecting all these steps we find the following solution in the future:

$$\varphi_{\mathbf{k}}(t) = \left(1 - \frac{i\lambda}{2\omega_{\mathbf{k}}}\right) \frac{e^{-i\omega_{\mathbf{k}}t}}{\sqrt{2\omega_{\mathbf{k}}}} + \frac{i\lambda}{2\omega_{\mathbf{k}}} \frac{e^{+i\omega_{\mathbf{k}}t}}{\sqrt{2\omega_{\mathbf{k}}}}$$

Read off the negative frequency components which correspond to out-particles, $|\beta_{\omega_{\mathbf{k}}}|^2 = \frac{\lambda^2}{4\omega_{\mathbf{k}}^2}$. This means that there is a non-zero particle number in the asymptotic future.

- So, as a recipe:
1. Construct exact solution in presence of potential via Lippmann–Schwinger.
 2. Make sure the free solution encodes the correct in-vacuum.
 3. Compute the late-time asymptotics.
 4. Read off the **Bogoliubov coefficient**.

But the result is not terribly exciting and could be guessed via dimensional analysis.



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PART III

EXACTLY SOLVABLE NON-LOCAL MODEL

non-local Green function ■ resonant particle creation

Non-local model

Let's consider instead this **non-local** model:

$$\exp \left[\ell^4 (\partial_t^2 + \omega_{\mathbf{k}}^2)^2 \right] (\partial_t^2 + \omega_{\mathbf{k}}^2) \varphi_{\mathbf{k}}(t) = -V(t) \varphi_{\mathbf{k}}(t)$$

The non-local retarded Green function satisfies

$$\exp \left[\ell^4 (\partial_t^2 + \omega_{\mathbf{k}}^2)^2 \right] (\partial_t^2 + \omega_{\mathbf{k}}^2) \mathcal{G}_{\mathbf{k}}^{\text{R}}(t' - t) = -\delta(t' - t) .$$

It has the following important property (DeWitt's “asymptotic causality”):

$$\mathcal{G}_{\mathbf{k}}^{\text{R}}(t' - t \pm \infty) = G_{\mathbf{k}}^{\text{R}}(t' - t)$$

This means that at early and late times physics is unchanged. **Or is it?**

Non-local model: surprising result!

Performing the same calculations again we find for the particle creation rate:

$$\beta_{\omega_{\mathbf{k}}} = \frac{i\lambda}{2\omega_{\mathbf{k}}} \frac{1}{1 - \lambda \mathcal{G}_{\mathbf{k}}^{\text{R}}(0)}$$

Observation: there is a new effect solely due to non-locality!

$$\mathcal{G}_{\mathbf{k}}^{\text{R}}(0) = \frac{\Gamma\left(\frac{3}{4}\right)\ell}{\pi} {}_2F_2\left(\frac{1}{4}, \frac{3}{4}; \frac{1}{2}, \frac{5}{4}; -k^4\ell^4\right) - \frac{\sqrt{2}k^2\ell^3}{6\Gamma\left(\frac{3}{4}\right)} {}_2F_2\left(\frac{3}{4}, \frac{5}{4}; \frac{3}{2}, \frac{7}{4}; -k^4\ell^4\right) \neq 0$$

This means that there exists a critical wave number for which the **particle creation rate diverges!**
(Provided the potential is positive and above a critical threshold: $\lambda\ell > \lambda_{\star}\ell = \pi/\Gamma\left(\frac{3}{4}\right) \approx 2.46369\dots$)

This effect is quite universal in this class of non-local theories, since it only assumes asymptotic causality. The behavior of the non-local Green function at $t = 0$ dictates the creation rate.

PART IV

CONCLUSIONS AND OUTLOOK

congratulations ■ we ■ made ■ it

Conclusions and outlook

We found an **unexpected resonant particle creation** due to non-locality.

- This effect is “non-perturbative” since it is an exact solution.
- It arises solely under the assumption of asymptotic causality.
- The scale of non-locality $\ell > 0$ introduces some sort of resonance since it smears out sharp delta-shaped objects.

Open questions/future directions:

- Does this happen for other potentials?
- Could it have implications for cosmology?
- What do you think?

Thank you for your attention :)



Abstract

Unexpected features of non-locality: resonant particle production

Let's consider a linear scalar field theory in the presence of an impulsive potential $\delta(t)$, which is an exactly solvable model. If there is a quantum mechanical vacuum at early times, then the potential term at $t=0$ creates a non-zero particle number at late times far into future. This means that the vacuum state of early times is mapped into a non-vacuum state at late times, and this can be described by so-called Bogoliubov coefficients in the framework of second quantization/Fock space quantization. In this talk I will extend these studies to an exactly solvable non-local model and explain how the future particle spectrum is impacted by the presence of non-locality. Surprisingly, there appears a strong resonant amplification of certain modes, leading to a burst of particles at late times.

Based on: Jens Boos, Valeri P. Frolov, and Andrei Zelnikov, “Resonant particle creation by a time-dependent potential in a nonlocal theory,” arXiv:2011.12929 [hep-th], submitted to Physics Letters B.