# Classical aspects of Poincaré gauge theory of gravity



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## Outline

- Motivation
- Gauging the Poincaré group
- Riemann–Cartan geometry
- Example: Einstein–Cartan theory
- Deformed Lie algebra

Based on F. W. Hehl, P. von der Heyde, G. D. Kerlick and J. M. Nester,
 "General Relativity with Spin and Torsion: Foundations and Prospects,"
 Rev. Mod. Phys. 48 (1976) 393.

History:

- field equations of General Relativity: Einstein (1915)
  - discovery of spin: Uhlenbeck & Goudsmit (1925)
  - description of spin: Pauli (1927), relativistic: Dirac (1928)
  - gauge theories: Weyl (1918, 1929, 1950), Yang–Mills (1954)
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The Dirac equation, minimally coupled to gravity (Weyl 1929):

 $i\gamma^{\alpha}e^{j}_{\alpha}\left(\partial_{j}+\frac{i}{4}\Gamma_{j}\right)\Psi+m\Psi=0$ 

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**What if** spin had been discovered before General Relativity? Would Einstein have applied the equivalence principle to a neutron instead?

Electrodynamics

- is invariant under the transformation  $A_i \rightarrow A_i + \partial_i \chi$
- couples to a conserved current j<sub>i</sub>

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#### $\rightarrow$ which gauge transformations on a matter field give rise to gravity?

Several possible answers: diffeomorphisms, translations, Lorentz rotations, Poincaré transformations, affine transformations, supersymmetric extensions, ...

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Why the Poincaré group?

- is the symmetry group of Minkowski space
- knows about energy-momentum and orbital angular momentum
- allows coupling to spin angular momentum (microphysics, cf. Synge)
- mass-spin classification (Wigner 1939)

## The Poincaré group

Poincaré group = {n translations}  $\rtimes$  { $\frac{n(n-1)}{2}$  Lorentz transformations}

Lie algebra:  $\begin{bmatrix} M_{\mu\nu}, M_{\alpha\beta} \end{bmatrix} = g_{\alpha[\mu} M_{\nu]\beta} - g_{\beta[\mu} M_{\nu]\alpha}, \quad \text{semidirect product} \\ \begin{bmatrix} M_{\mu\nu}, P_{\alpha} \end{bmatrix} = g_{\alpha[\mu} P_{\nu]}, \\ \begin{bmatrix} P_{\mu}, P_{\alpha} \end{bmatrix} = 0.$ 

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Conserved Noether currents:

energy-momentum: 
$$\partial_{i}\mathfrak{T}_{k}^{i} = 0$$
  
 $\mathfrak{T}_{k}^{i} \coloneqq \mathcal{L}\delta_{k}^{i} - \frac{\partial \mathcal{L}}{\partial(\partial_{i}\Phi^{A})}(\partial_{k}\Phi^{A})$   
total angular momentum:  $\partial_{i}(\mathfrak{S}_{kl}^{i} + \mathbf{x}_{[k}\mathfrak{T}_{l]}^{i}) = 0$   
 $\mathfrak{S}_{kl}^{i} \coloneqq \frac{\partial \mathcal{L}}{\partial(\partial_{i}\Phi^{A})}f_{[kl]}^{A}{}_{B}\Phi^{B}$ 

Note: the semidirect product structure is everywhere.

Starting point: a field theory in Minkowski space,  $\mathcal{L} = \mathcal{L}(\phi^A, \partial \phi^A)$ 

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Action on the field (generators satisfy Poincaré algebra):

$$\phi^{\mathsf{A}} \mapsto \Pi \phi^{\mathsf{A}} \coloneqq \left( \delta^{\mathsf{A}}_{\mathsf{J}} + \Lambda^{\alpha\beta} \mathsf{f}_{\alpha\beta}{}^{\mathsf{A}}_{\mathsf{J}} - \delta^{\mathsf{A}}_{\mathsf{J}} \Lambda_{\alpha}{}^{\beta} \delta^{\alpha}_{\mathsf{i}} \mathsf{x}^{\mathsf{i}} \partial_{\beta} - \delta^{\mathsf{A}}_{\mathsf{J}} \epsilon^{\alpha} \partial_{\alpha} \right) \phi^{\mathsf{J}}(\mathsf{x})$$

rotation induced translation  $\rightarrow$  angular momentum

Reminder: 
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Invariance is spoiled:  $\delta S = (\Pi S) - S = \int \left[ \left( -\partial_i \Lambda^{\alpha\beta} \right) \mathfrak{S}_{\beta\alpha}{}^i + \left( \partial_i \epsilon^{\gamma} - \Lambda_{\beta}{}^{\gamma} \delta_i^{\beta} \right) \mathfrak{T}_{\gamma}{}^i \right] d^4x$ 

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How to rescue invariance? Introduce gauge potentials:

$$\begin{split} \delta^{i}_{\mu} &\mapsto e^{i}_{\mu}(x), &= \text{tetrad, 4 translational gauge potentials} \\ \partial_{i} &\mapsto D_{i} \coloneqq \partial_{i} - \Gamma_{i}^{\alpha\beta}(x) f_{\alpha\beta} &= \text{connection, 6 rotational gauge potentials} \end{split}$$

Inhomogeneous transformation of gauge potentials ensures local Poincaré invariance.

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Gauge field strengths:

$$\begin{split} \mathsf{F}_{\mathbf{ij}}^{\ \mu} &\coloneqq 2\left(\partial_{\left[\mathbf{i}^{\ \mathbf{e}_{\mathbf{j}}\right]}^{\ \mu}} + \mathsf{\Gamma}_{\left[\mathbf{i}^{\ \mu\alpha} \mathbf{e}_{\mathbf{j}}\right]}^{\ \beta} \mathsf{g}_{\alpha\beta}\right) &= \mathsf{translational\ curvature} \\ \mathsf{F}_{\mathbf{ij}}^{\ \mu\nu} &\coloneqq 2\left(\partial_{\left[\mathbf{i}^{\ \mathbf{f}_{\mathbf{j}}\right]}^{\ \mu\nu}} + \mathsf{\Gamma}_{\left[\mathbf{i}^{\ \alpha\mu} \mathsf{\Gamma}_{\mathbf{j}}\right]}^{\ \beta\nu} \mathsf{g}_{\alpha\beta}\right) &= \mathsf{rotational\ curvature} \end{split}$$

## Recognizing Riemann–Cartan geometry

The field strengths  $F_{ij}^{\mu}$  and  $F_{ij}^{\mu\nu}$  as well as the covariant derivative  $D_i$  allow for a geometric interpretation of the resulting structure: a Riemann–Cartan geometry.

 $\begin{array}{rcl} {\sf F}_{ij}{}^{\mu} &\coloneqq {\sf T}_{ij}{}^{\mu} &= {\sf torsion} \\ {\sf F}_{ij}{}^{\mu\nu} &\coloneqq {\sf R}_{ij}{}^{\mu\nu} &= {\sf Riemann-Cartan\ curvature} \end{array}$ 

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Geometric description in terms of tensor-valued differential forms (Trautman 1973):

$$\vartheta^{\mu} = e_{a}{}^{\mu}dx^{i} \qquad \qquad \mathsf{T}^{\mu} \coloneqq \frac{1}{2}\mathsf{T}_{ij}{}^{\mu}dx^{i} \wedge dx^{j} = d\vartheta^{\mu} + \mathsf{\Gamma}^{\mu}{}_{\alpha} \wedge \vartheta^{\alpha} \qquad \qquad \mathsf{D}\mathsf{T}^{\mu} = \mathsf{R}^{\mu}{}_{\alpha} \wedge \vartheta^{\alpha}$$

$$\Gamma^{\mu\nu} = \Gamma_{i}^{\mu\nu} dx^{i} \qquad \qquad R^{\mu}_{\nu} \coloneqq \frac{1}{2} R_{ij}^{\mu}_{\nu} dx^{i} \wedge dx^{j} = d\Gamma^{\mu}_{\nu} + \Gamma^{\mu}_{\alpha} \wedge \Gamma^{\alpha}_{\nu} \qquad DR^{\mu}_{\nu} = 0$$

A more direct interpretation is possible in terms of holonomy.

#### Geometric interpretation of curvature and torsion

Parallel transport a vector along A  $\rightarrow$  B  $\rightarrow$  C  $\rightarrow$  D  $\rightarrow$  E  $\rightarrow$  A:



The holonomy of parallel transport is a Poincaré transformation:  $\Delta v^{\mu} = A^{\mu}{}_{\alpha}v^{\alpha} \oplus b^{\mu}$ 

## Example: Einstein–Cartan theory (1/2)

The simplest extension of Einstein's General Relativity (Sciama 1960, Kibble 1961).

$$\begin{split} S_{\text{EC}} &= \frac{1}{2\kappa} \int \left( \mathsf{R}^{\alpha\beta} \wedge \eta_{\alpha\beta} - 2\Lambda\eta \right) + \mathcal{L}_{\text{mat}} \\ \frac{\delta \mathsf{S}_{\text{EC}}}{\delta \vartheta^{\mu}} &= 0 \quad \Rightarrow \quad \frac{1}{2} \eta_{\mu\alpha\beta} \wedge \mathsf{R}^{\alpha\beta} + \Lambda \eta_{\mu} = \kappa \mathfrak{T}_{\mu} \\ \frac{\delta \mathsf{S}_{\text{EC}}}{\delta \Gamma^{\mu\nu}} &= 0 \quad \Rightarrow \quad \frac{1}{2} \eta_{\mu\nu\alpha} \wedge \mathsf{T}^{\alpha} = \kappa \mathfrak{S}_{\mu\nu} \end{split}$$

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The source of curvature is the canonical energy-momentum  $\mathfrak{T}_{\mu} = \delta \mathcal{L}_{mat} / \delta \vartheta^{\mu}$ , and torsion is linked algebraically to the spin angular momentum  $\mathfrak{S}_{\mu\nu} = \delta \mathcal{L}_{mat} / \delta \Gamma^{\mu\nu}$ .

Note that energy-momentum and spin angular momentum are no longer conserved.

## Example: Einstein–Cartan theory (2/2)

Rewrite field equations: Substitute spin for torsion and split off post-Riemannian part. Relate to Einstein equations by splitting energy momentum into symmetric energymomentum  $\Theta_{ij} = \Theta_{ji} = \delta \mathcal{L}_{mat} / \delta g^{ij}$  and find  $\mathfrak{T}_{ij} = \Theta_{ij} + (\kappa \mathfrak{S}^2)_{ij}$ .

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When does spin-spin interaction balance the gravitational collapse?

mass density of spins  $\approx \kappa$  (spin density)<sup>2</sup>

 $\rightarrow r_{EC} \approx \left(\lambda_{Compton} \times \ell_{Planck}^2\right)^{1/3} \approx 10^{-26} \text{cm} \gg \ell_{Planck} \approx 10^{-33} \text{cm}$ 

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Applications:

- contact interaction affects energy-momentum and hence singularity theorems
- bouncing cosmologies (Poplawski 2012; Magueijo, Zlosnik, Kibble 2013; …)

## The deformed Lie algebra in Poincaré gauge theory (1/2)

Poincaré gauge theory is an external gauge theory, unlike Yang–Mills. Therefore, the localized symmetries affect the Lie algebra of the gauge group.

$$\begin{bmatrix} \mathsf{M}_{\mu\nu}, \mathsf{M}_{\alpha\beta} \end{bmatrix} = \mathsf{g}_{\alpha[\mu}\mathsf{M}_{\nu]\beta} - \mathsf{g}_{\beta[\mu}\mathsf{M}_{\nu]\alpha}, \\ \begin{bmatrix} \mathsf{M}_{\mu\nu}, \mathsf{D}_{\alpha} \end{bmatrix} = \mathsf{g}_{\alpha[\mu}\mathsf{D}_{\nu]}, \\ \begin{bmatrix} \mathsf{D}_{\mu}, \mathsf{D}_{\alpha} \end{bmatrix} = \mathsf{F}_{\mu\alpha}{}^{\rho\sigma}\mathsf{M}_{\rho\sigma} - \mathsf{F}_{\mu\alpha}{}^{\sigma}\mathsf{D}_{\sigma} \neq \mathbf{0}$$

Is this algebra closed?

$$\begin{pmatrix} 2 \\ \Lambda(x), \hat{\epsilon}(x) \end{pmatrix} \circ \begin{pmatrix} 1 \\ \Lambda(x), \hat{\epsilon}(x) \end{pmatrix} \stackrel{?}{=} \begin{pmatrix} 3 \\ \Lambda(x), \hat{\epsilon}(x) \end{pmatrix}$$

Yes, see Hehl (1979). Closure relations:

$$\begin{array}{rcl} \overset{3}{\epsilon}{}^{\alpha} &=& -\overset{2}{\Lambda}{}_{\gamma}{}^{\alpha}\overset{1}{\epsilon}{}^{\gamma} + \overset{1}{\Lambda}{}_{\gamma}{}^{\alpha}\overset{2}{\epsilon}{}^{\gamma} - \overset{2}{\epsilon}{}^{\beta}\overset{1}{\epsilon}{}^{\gamma}\mathsf{F}_{\beta\gamma}{}^{\alpha}, \\ \overset{3}{\Lambda}{}_{\alpha}{}^{\beta} &=& -\overset{2}{\Lambda}{}_{\alpha}{}^{\gamma}\overset{1}{\Lambda}{}_{\gamma}{}^{\beta} + \overset{1}{\Lambda}{}_{\alpha}{}^{\gamma}\overset{2}{\Lambda}{}_{\gamma}{}^{\beta} - \overset{2}{\epsilon}{}^{\gamma}\overset{1}{\epsilon}{}^{\delta}\mathsf{F}_{\gamma\delta\alpha}{}^{\beta} \end{array}$$

## The deformed Lie algebra in Poincaré gauge theory (2/2)

Poincaré gauge theory is an external gauge theory, unlike Yang–Mills. Therefore, the localized symmetries affect the Lie algebra of the gauge group.

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Open questions:

- Is it still possible to perform a mass-spin classification à la Wigner (1939)?
- Is there a relation to the deformed Poincaré algebra in relative locality?

Thank you for your attention.

#### References

 F. W. Hehl, P. von der Heyde, G. D. Kerlick and J. M. Nester,
 "General Relativity with Spin and Torsion: Foundations and Prospects," Rev. Mod. Phys. 48 (1976) 393.

[2] F. W. Hehl, "Four Lectures on Poincaré Gauge Field Theory," Erice, 1979.

## Classical aspects of Poincaré gauge theory of gravity

I will briefly highlight a few cornerstones in the development of gauge theory, and then proceed to the gauge structure present in gravity. Following [1], I will argue that if one wishes to take the fermionic character of matter into account, the Poincaré group will give rise to a satisfactory gauge-theoretical description of gravity. This will include both energy-momentum and spin-angular momentum as sources of the gravitational field.

In a second step, I will elaborate on the emerging structure of a Riemann–Cartan geometry. Einstein–Cartan theory will be sketched, a minimal and viable gauge-theoretical extension of Einstein's General Relativity. If time permits, I will briefly mention its implications for cosmology and the possible resolution of singularities.

I will close by pointing out the deformed Lie algebra of the Poincare group as a result of the gauging procedure: unlike in Yang–Mills theory with its internal symmetry groups, here the Lie algebra is deformed due to the presence of curvature and torsion. The implications of this deformation, both in the classical and quantum regime, remain to be seen.